Pg, the real deal*

This force holds object in (approximately) circular orbits.

$$
1 S S=g \pi 9 \frac{\pi}{k_{g}}
$$

g is called: the GRAVITATIONAL FIELD STRENGTH a gravitational field is the bending of space-time by a mass.

$$
\begin{aligned}
& \vec{g}^{s}=\frac{G n}{r^{2}} \text { on Earth }=9.8 \frac{\mathrm{~N}}{\mathrm{~kg}} \\
& \vec{J}_{\text {min }}=6.67 \times 10^{-11}\left(\frac{7.35 \times 10^{22}}{(1.74 \times 106)^{2}}\right)=1.62 \frac{\mathrm{~N}}{\mathrm{~kg}}
\end{aligned}
$$

$$
\text { Find Eg Earth } \rightarrow \text { moon } 20 \quad \text { radius } \in \rightarrow M
$$

$$
\begin{aligned}
& \text { Earth } \rightarrow \text { moon radius } \quad=3.84 \times 10^{8} \mathrm{~m} \\
& =1.98 \times 10^{20} \mathrm{~N}
\end{aligned}
$$

$A$

$$
=\frac{G m_{1} r_{2}}{r^{2}}=
$$

The moon orbits the Earth, we can use this To find the mass of the Earth

geo $\rightarrow$ Earth


Syn $\rightarrow$ Sane chronous $\rightarrow$ Time

Prove there is ONLY one radial distance from the Earth which allows for geosynchronous orbit.

$$
\begin{aligned}
& \text { Ff exists between any two masses and is given by: } \\
& F_{s}=\frac{G_{m_{1} m_{2}}^{2}}{r^{2}} \quad G=6.67 \times 10^{-11} \mathrm{~N} \frac{\mathrm{n}^{2}}{\mathrm{k}_{\mathrm{s}}^{2}}
\end{aligned}
$$

Prove there is ONLY one radial distance from the Earth which allows for geosynchronous orbit.

$$
\begin{aligned}
& 4 \pi^{2} r f^{2}=\frac{G m_{1} m_{2}}{r^{2}} \\
& \frac{4 \pi^{2} r^{3}}{G m_{1} m_{2}}=T^{2}
\end{aligned}
$$

In orbit there is no surface that objects rest upon, this means that $\mathrm{Fn}=\mathrm{ON}$

GEOSYNCHRONOUS (geostationary)


Apollo. 13
When in orbit an object is constantly
Accelerating toward the middle of circle, it is "freely falling" at $a_{c}=g$
effective

When free falling objects have apparent weight (Fn)=0 VOMit Comet
What is the apparent weight of an astronaut in orbit in the ISS?
$\vec{g}$ of Jupiter is $\approx 25 \frac{\mathrm{~N}}{\mathrm{~kg}}$
Find the gravitational field $(\vec{y})$ at $4 x$

$$
\begin{array}{ll}
\text { Jupiter radius. } & r_{\text {Jupiter }}=70 \mathrm{Mm} . \\
& 70,000,000 \mathrm{~m}
\end{array}
$$

$$
\begin{aligned}
& \vec{g}=\frac{G m}{r^{2}} \\
& 25=\frac{6.67 \times 10^{-11} m \quad \text { decrease by } \frac{1}{16} \rightarrow \frac{1}{4^{2}}}{4 r)^{2}}
\end{aligned}
$$

Find the gravitational field strength at a HEIGHT of $3.0 \times 10^{6} \mathrm{~m}$ above Earth's surface.

$$
\begin{aligned}
\vec{S}=\frac{G_{m}}{r^{2}} & =\frac{6.67 \times 10^{-11}\left(5.98 \times 10^{24}\right)}{\left(6.38 \times 10^{6}+3.0 \times 10^{6}\right)^{2}} \\
& =\frac{4.53 \frac{\mathrm{~N}}{\mathrm{~kg}} G_{r}}{G g}
\end{aligned}
$$

Determine the orbital velocity at that HEIGHT!!!

$$
\begin{array}{rlrl}
F_{c} & =F_{y} & \\
v a_{c} & =m y & V & =6.52 \times 10^{3} \mathrm{~m} / \mathrm{s} \\
\frac{v^{2}}{r} & =g & & =6.5 \mathrm{kr} / \mathrm{s}
\end{array}
$$

Inverse Square Law



$$
\begin{aligned}
\vec{g} @ \text { Earth's Surface } & =9.8 \\
& =\frac{G m}{r^{2}}
\end{aligned}
$$

Find $\vec{g} @ 2$ Earth radii
Thomas:
"I could be the first man to never come at if a black hole."

Pentose
from center.

$$
\frac{G m}{(2 r)^{2}}
$$

3 Earth radii

$$
\begin{array}{r}
\vec{g}=\frac{G m}{(3 r)^{2}} \\
\quad 49
\end{array}
$$


"Beyond your. imagination."

imagination.

$$
\begin{aligned}
& \text { Big } \text { Sun. } \\
& \text { Fast } \approx 10 \% \\
& \text { An exoplanet has gravitational field strength } \\
& \text { Of } 36 \mathrm{~N} / \mathrm{kg} \text { at its surface, what is } \mathrm{g} \text { at a HEIGHT of } \\
& 5 \text { radii }
\end{aligned}
$$

## \&

